16 Jun 2004, NBI, Copenhagen

Peter Skands (Lund)

(report on work performed with W. Porod, ITP U. Zürich)

Measuring Neutrino Mixing at LHC?

— How to measure neutrino mixing using only a hadron collider and no water.



See also:

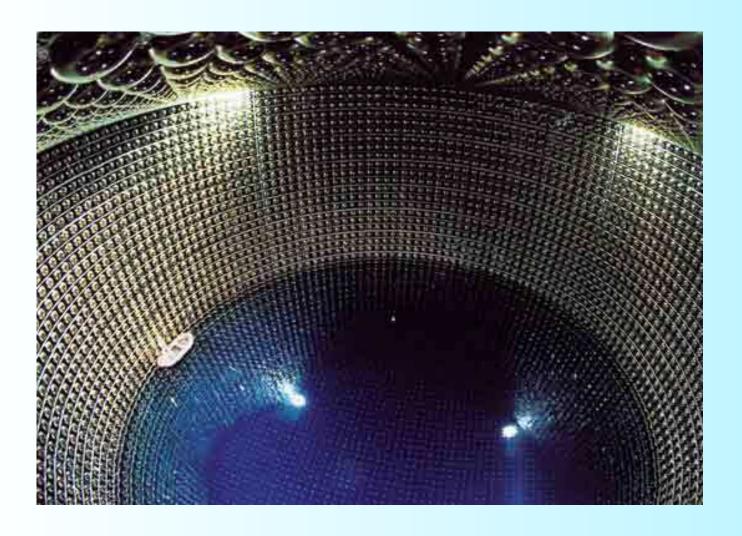
W. Porod+PS, hep-ph/0401077

M. Hirsch+W. Porod, Phys.Rev.D68:115007(2003)

W. Porod+M. Hirsch+J. Romao+J. Valle, Phys.Rev.D63:115004(2001)

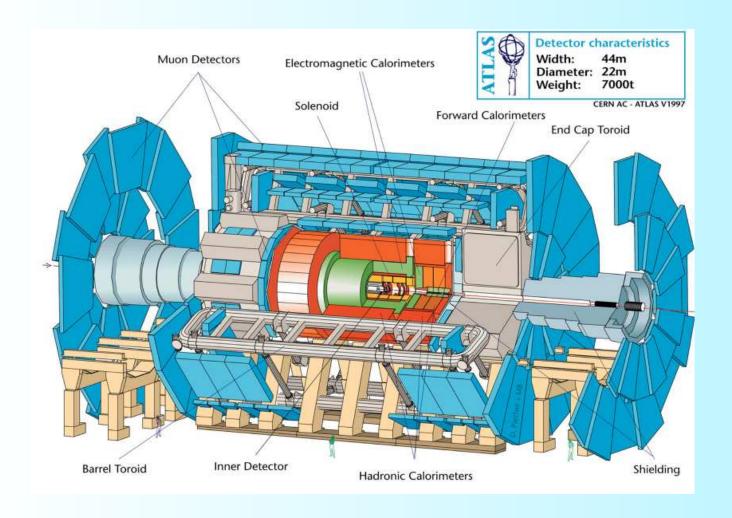
Sanity Check

This is a detector for neutrino physics.



Sanity Check

This is not a detector for neutrino physics.



Overview

- 1. Fast Forward SUSY Intro.
- 2. R-Parity and R-Parity Violation.
- 3. R–Parity Violation with Bilinear Terms: a SUSY origin of ν masses?
- 4. Measuring a ν angle at a hadron collider?

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Supersymmetry.



Presumably more about this in training course tomorrow! In fast–forward mode, supersymmetry is:

- The only (interesting) way of extending the known space—time symmetries.
- A fundamental relation between fermionic and bosonic degrees of freedom.
- The "super" in superstrings.

TeV-Scale SUSY is popular among BSM theories because:

- It provides an elegant solution to the hierarchy problem.
- It predicts a natural dark matter candidate.
- It paves the way for natural grand unification.
- It is something experimentalists can look for.

Supersymmetry.



SUPERSYMMETRY

For every boson, there is a fermion For every fermion, there is a boson

 $S = \frac{1}{2}$

photon +
$$W^{\pm}$$
 and Z^{0} + gluon

Supersymmetry.



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For every boson, there is a fermion For every fermion, there is a boson

photon +
$$W^{\pm}$$
 and Z^{0} + gluon

$$S = \frac{1}{2}$$

$$S=0$$

$$S = \frac{1}{2}$$

$$S = \frac{1}{2}$$

+ (at least) another Higgs doublet.

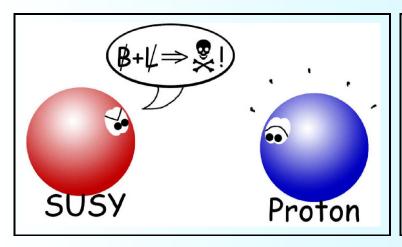
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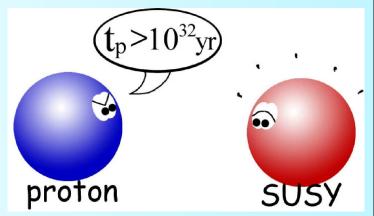
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Aberdabei

So SUSY is really popular, but...

in *any* Supersymmetric extension of the Standard Model (SM), the first problem you have to deal with, is how to avoid **rapid proton decay!**





BNV+LNV together is a bad cocktail!

ullet Write down all (renormalizable) terms consistent with SM gauge invariance and (N=1) Supersymmetry ullet

$$W_{\text{SUSY}} = W_{\text{MSSM}} + W_{\text{BNV}} + W_{\text{LNV}}$$

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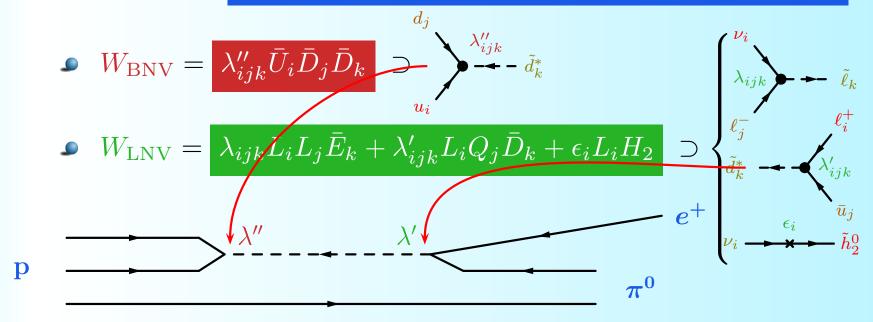
$$W_{\rm BNV} = \lambda_{ijk}^{\prime\prime} \bar{U}_i \bar{D}_j \bar{D}_k$$

$$W_{\text{LNV}} = \lambda_{ijk} L_i L_j \bar{E}_k + \lambda'_{ijk} L_i Q_j \bar{D}_k + \epsilon_i L_i H_2$$

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$$W_{\text{SUSY}} = W_{\text{MSSM}} + W_{\text{BNV}} + W_{\text{LNV}}$$



ullet fast proton decay $\propto rac{|\lambda'||\lambda''|}{M_{
m SUSY}^2} \stackrel{M_{
m SUSY}}{\Longrightarrow}^{1{
m TeV}} |\lambda'||\lambda''| \lesssim 10^{-30}$.

What can be done?

- Start with naive, simple guesses:
 - 1. Lepton Number is conserved?
 - 2. Baryon Number is conserved?
 - 3. Both are conserved (R-parity = $(-1)^{L+3B+2S}$)?

What can be done?

- Start with naive, simple guesses:
 - 1. Lepton Number is conserved?
 - 2. Baryon Number is conserved?
 - 3. Both are conserved (R-parity = $(-1)^{L+3B+2S}$)?
- ightharpoonup R-parity has interesting consequences:

SM particles have R=+1 and SUSY ones R=-1

- → sparticles are born and killed by the pair ... so sparticles from Big Bang decayed to "LSP" are still around?
- → solution to dark matter problem in cosmology?
- But no deep theoretical motivation...

Simple guesses ain't so bad!

What can we say about the symmetry we're looking for?

- Only discrete gauge symmetries are absolutely stable.
- But you can find discretized versions of gauge symmetries which reduce to both R-parity as well as symmetries equivalent to Baryon and Lepton number conservation.
- Again no clear preference for one or the other.

What about Beyond-the-MSSM contributions?

- Anyway, all this was just to make the case...

All possibilities should be considered!

R conservation vs R violation.

$$W_{\text{SUSY}} = W_{\text{MSSM}}(+W_{\text{BNV}} + W_{\text{LNV}})$$

- - Conserves $R \rightarrow$ only allows sparticles to be produced in pairs and does not mediate LSP decay.
- Signature is missing (transverse) energy from escaping LSP's.
- At the LHC, squark and gluino pair production will dominate over most of parameter space.
- Typically, squarks and gluinos are among the heavier sparticles, hence other typical features are multiple jets and/or leptons which are split off in a chain of decays to lighter and lighter sparticles, ending with the (stable) LSP.
- Escaping LSP → tricky mass reconstruction (use edges).

Trilinear Lepton Number Violation

$$W_{\text{SUSY}} = W_{\text{MSSM}} + W_{\text{LNV}}$$

- - "LLE" ("LQD") allows single slepton production at a linear collider (hadron collider). "LQD" also allows resonant squark/slepton production at an ep collider.
- Rich phenomenology. With just 2- and 3-body decays of sparticles to particles, more than 1200 new decay chanels!

$$\mathsf{LLE}(\lambda) \colon \ \, \bullet \ \, \tilde{e}_j^- \to \bar{\nu}_i \ell_k^-, \; \nu_k \ell_i^- \qquad \mathsf{LQD}(\lambda') \colon \ \, \bullet \ \, \tilde{e}_i^- \to \bar{u}_j d_k$$

HERWIG: P.Richardson,hep-ph/0101105
PYTHIA: PS,Eur.Phys.J.C23:173(2002)

$$ilde{
u}_i
ightarrow ar{d}_i d_k$$

$$\sim$$
 $+1$

Trilinear Baryon Number Violation

$$W_{\rm SUSY} = W_{\rm MSSM} + W_{\rm BNV}$$

"UDD", violates Baryon Number. Allows single (resonant) squark production from qq initial state. Allows 2-body decays of squarks to quarks and 3-body decays of gauginos to quarks.

UDD
$$(\lambda'')$$
:

$$\bullet$$
 $\tilde{u}_i \rightarrow \bar{d}_i \bar{d}_k$

HERWIG: P. Richardson, hep-ph/0101105

PYTHIA: T. Sjöstrand+PS, Nucl.Phys.B659:243(2003)

NB: Unique colour structures → new colour topologies not adressed by standard fragmentation schemes. Detailed dynamical modelling so far developed only for string fragmentation (implemented in PYTHIA).

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Neutrino Summary





"I have done a terrible thing, I have invented a particle that cannot be detected"

W. Pauli



Masatoshi Koshiba

Raymond Davis Jr.





Nobel prize 2002: Neutrinos have mass!

Neutrino Summary

Neutrino sector: a window to physics beyond SM?

- 1. Too few ν_{μ} from atmosphere, can be explained by oscillations into ν_{τ} : $\Delta m^2_{\rm atm} = m_3^2 m_2^2 \sim 10^{-3} 10^{-2}\,{\rm eV}^2$
- 2. Too few $\nu_{\rm e}$ from Sun, can be explained by oscillations into ν_{μ} : $\Delta m_{\rm sol}^2 = m_2^2 m_1^2 \sim 10^{-5} 10^{-4}\,{\rm eV}^2$
- 3. Bi-maximal mixing pattern: θ_{23} large, θ_{12} large, and θ_{13} small.

Explanations generally look like this:

$$\begin{pmatrix} 0 & m \\ m & M \end{pmatrix}$$

Neutrino Masses

$$\begin{pmatrix} 0 & m \\ m & M \end{pmatrix}$$
 m : Dirac mass. Electroweak scale? M : Majorana mass. High scale?

$$\lambda_{\pm} = \frac{1}{2}M^2 \pm \frac{1}{2}\sqrt{M^2 + 4m^2} \sim \begin{cases} -m^2/M = m_{\nu} \\ M + m^2/M \gtrsim 10^9 \, \mathrm{GeV} \end{cases}$$

→ No fun for high energy colliders...

An alternative possibility could be M of order M_Z , with m thus being rather small...

Bilinear R-violation

$$W_{\rm SUSY} = W_{\rm MSSM} + \epsilon_i L_i H_2$$

(Occurs e.g. when R-parity is broken spontaneously)

For us, the important consequences are:

EW symmetry is broken by Higgs and sneutrino vev's,

$$\langle \nu_i \rangle = v_i$$
 (i.e. $m_W^2 = \frac{1}{4} g^2 (v_d^2 + v_u^2 + v_1^2 + v_2^2 + v_3^2)$).

▶ Neutrinos mix with neutralinos \rightarrow 7 × 7 mixing:

In block form:
$$M_N = \begin{pmatrix} 0 & m_{(3\times4)} \\ m^T_{(4\times3)} & M_{(4\times4)} \end{pmatrix}$$

Bilinear R-Violation

Determining the masses:

- Find diagonalizing matrix: $N^*M_NN^{-1} = \operatorname{diag}(m_{\nu_i}, m_{\tilde{\chi}_j^0})$.
- ullet First transform M_N to block-diagonal:

$$N^*M_NN^{-1}= ilde{N}^*\left(egin{array}{cc} m_{
m eff} & 0 \ 0 & M_{ ilde{\chi}^0} \end{array}
ight) ilde{N}^{-1}$$
 ; $ilde{N}=\left(egin{array}{cc} V_{
u}^{\dagger} & 0 \ 0 & N_{ ilde{\chi}^0} \end{array}
ight)$

• The matrix m_{eff} is projective, looks like:

$$m_{ ext{eff}} = rac{M_1 g^2 + M_2 g'^2}{4 ext{det}(M_{ ilde{\chi}^0})} \left(egin{array}{ccc} \Lambda_e^2 & \Lambda_e \Lambda_{\mu} & \Lambda_e \Lambda_{ au} \\ \Lambda_e \Lambda_{\mu} & \Lambda_{\mu}^2 & \Lambda_{\mu} \Lambda_{ au} \\ \Lambda_e \Lambda_{ au} & \Lambda_{\mu} \Lambda_{ au} & \Lambda_{ au}^2 \end{array}
ight)$$

$$\Lambda_i = \mu v_i + v_d \epsilon_i$$

Bilinear R-Violation

So only 1 non-zero eigenvalue in m_{eff} !

- $N_{\tilde{\chi}^0}^* M_{\tilde{\chi}^0} N_{\tilde{\chi}^0}^{\dagger} = \operatorname{diag}(m_{\tilde{\chi}_i^0})$
- $V_{\nu}^T m_{\rm eff} V_{\nu} = {\rm diag}(0,0,m_{\nu})$; $m_{\nu} = {\rm Tr}(m_{\rm eff}) \propto \Lambda^i \Lambda_i$

1 neutrino becomes massive at tree level.

(Remaining neutrinos acquire mass at 1 loop).

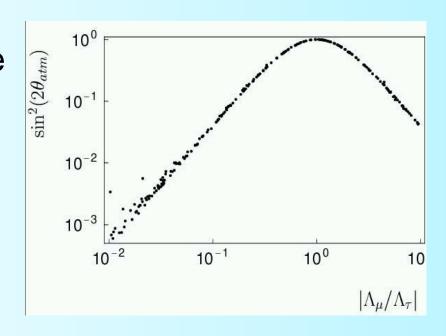
$$V_{\nu} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos \theta_{23} & -\sin \theta_{23} \\ 0 & \sin \theta_{23} & \cos \theta_{23} \end{pmatrix} \times \begin{pmatrix} \cos \theta_{13} & 0 & -\sin \theta_{13} \\ 0 & 1 & 0 \\ \sin \theta_{13} & 0 & \cos \theta_{13} \end{pmatrix}$$

• NOTE:
$$\hookrightarrow$$
 $\tan \theta_{13} = \frac{\Lambda_e}{\sqrt{\Lambda_\mu^2 + \Lambda_\tau^2}}$; $\tan \theta_{23} = -\frac{\Lambda_\mu}{\Lambda_\tau}$.

Bilinear R-Violation

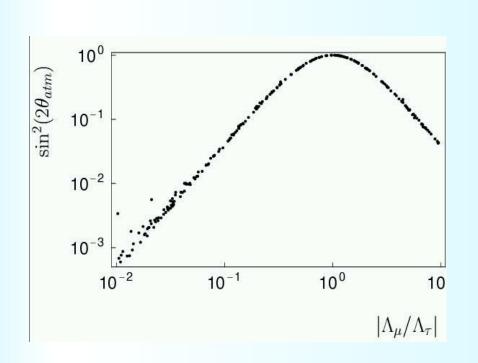
Details depend on the particular SUSY scenario, but the general results are:

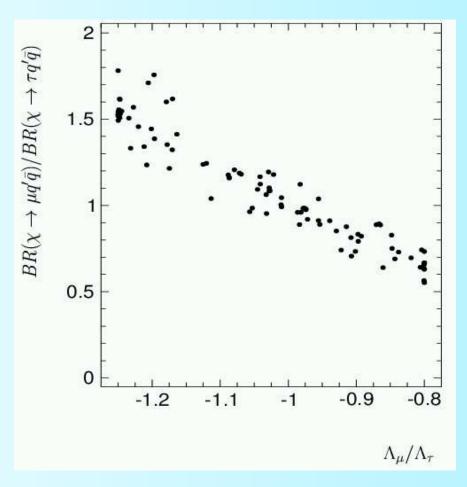
- The tree-level mass m_{ν} generates the atmospheric mass scale.
- The loop-induced (=small) corrections generate the Solar mass scale (→ hierarchical mass pattern).
- With $\Lambda_e \ll \Lambda_\mu \sim \Lambda_\tau$, the bi-maximal mixing can be accomodated.



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- BRPV couplings also responsible for LSP decay.
- ightharpoonup Ratio of $\tilde{\chi}_1^0$ semileptonic branching ratios is strongly correlated with $\Lambda_i/\Lambda_i!$

 \rightarrow model of SUSY origin of ν mass can be checked by "measuring" θ_{atm} at a hadron collider:

$$\tan^2 \theta_{\text{atm}} \simeq \left| \frac{\Lambda_{\mu}}{\Lambda_{\tau}} \right|^2 \simeq \frac{BR(\tilde{\chi}_1^0 \to \mu^{\pm} W^{\mp})}{BR(\tilde{\chi}_1^0 \to \tau^{\pm} W^{\mp})}$$

Note: this prediction is independent of the R-conserving MSSM parameters.

To illustrate method, we have investigated a specific example, based on the SPS1a mSUGRA point. (using SPHENO 2.2 together with PYTHIA 6.3, and SLHA interface to pass parameters)

The R-Violating parameters are (in MeV):

$$\epsilon_i = (43, 100, 10)$$
 $v_i = (-2.9, -6.7, -0.5)$

(chosen to fit neutrino data)

Total SUSY cross section for SPS1a: $\sigma_{\rm SUSY} \sim 41 {\rm pb}$.

Some shortcuts:

Detector resolution and hadronization effects are ignored

The "detector":

- Calorimeter: $|\eta| < 4.9$. Inner detector: $|\eta| < 2.5$.
- Electrons ($|\eta| < 2.5$, $p_{\perp} > 5 \, \text{GeV}$, $\varepsilon = 75\%$)
- ho Muons ($|\eta| < 2.5$, $p_{\perp} > 6 \, \text{GeV}$, $\varepsilon = 95\%$)
- ightharpoonup Taus ($|\eta| < 2.5$, $\mathrm{p}_{\perp} > 20\,\mathrm{GeV}$, $\varepsilon_{\mathrm{3-prong}} = 85\%$)
- ightharpoonup Vertex resolution: 20 μ transverse and 0.5mm longitudinal.
- "Triggers": 4j100, 2j100+e20/mu20, j100+2(e20/mu20).

Events characterized by:

- Since R-Violating parameters are small, the only real deviation from MSSM phenomenology is LSP decay.
- → pair production of squarks/gluinos dominate, with cascades down to LSP, which subsequently decays through LNV.
- Also due to smallness of LNV parameters, LSP is long–lived. Decay length here is $c\tau=0.5$ mm.
- → very clean signature: 2 reconstructed detached vertices in fair fraction of signal events (define vertex reconstruction ellipsoid = 5 times resolution and reject tracks that intersect it).

```
100 fb<sup>-1</sup> of data: \sim 10000 reconstructed \tilde{\chi}_1^0 \to \mu W \to \mu q \overline{q}' decays. \sim 1500 reconstructed \tilde{\chi}_1^0 \to \tau W \to \tau_{3-prong} q \overline{q}' decays.
```

 \rightarrow precision on $\tan^2 \theta_{\rm atm}$ is \sim couple of percent

Summary & Conclusion

- Parity: conserved or violated? Nobody knows...
- Possible sources of RPV: UDD, LLE, LQD, and Bilinear.
- Bilinear RPV has interesting consequences:

Sneutrinos play rôle of extra Higgses and acquire vevs.

Neutrinos mix with Neutralinos.

A "low-scale" seesaw mechanism results, whereby neutrinos become massive.

Models consistent with neutrino data give predictions which can be tested at hadron colliders.

So all we can really say is:

This may not seem to be a detector for neutrino physics...

